Hubble 26 years of utilization and on-orbit servicing



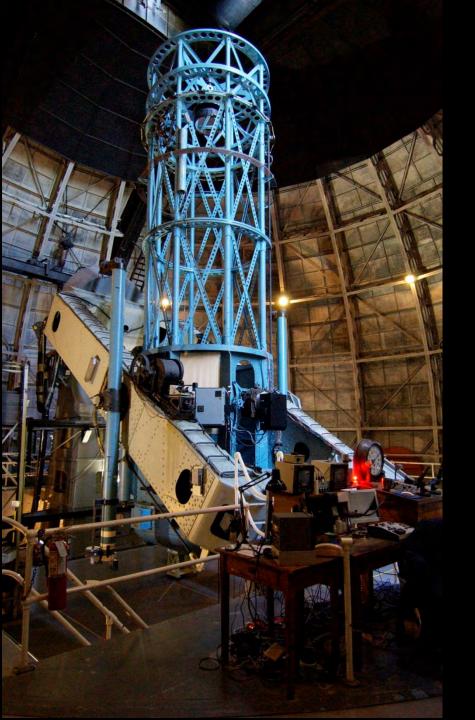


Hubble

the Astronomer

Edwin Powell Hubble, 1889-1953

University of Chicago (math, astronomy), Oxford University (law) US Army (WW1), Yerkes Observatory, Mt Wilson Observatory, Palomar Observatory



2.5 m Hooker Telescope Mt Wilson Observatory

Discoveries made using this instrument:

The Universe is much bigger than the Milky Way Galaxy!

Hubble Law (1922-1925)

Expansion of the Universe

Classification of Galaxies

Hubble's Law

The dominant motion in the universe is a smooth expansion expressed by the **Hubble's Law**:

Recessional Velocity of galaxies = Hubble Constant times their distance

$$V = H_0 D$$

where

V is the observed velocity of the galaxy away from us, in km/sec H₀ is Hubble's "constant", in km/sec/Mpc D is the distance to the galaxy in Mpc

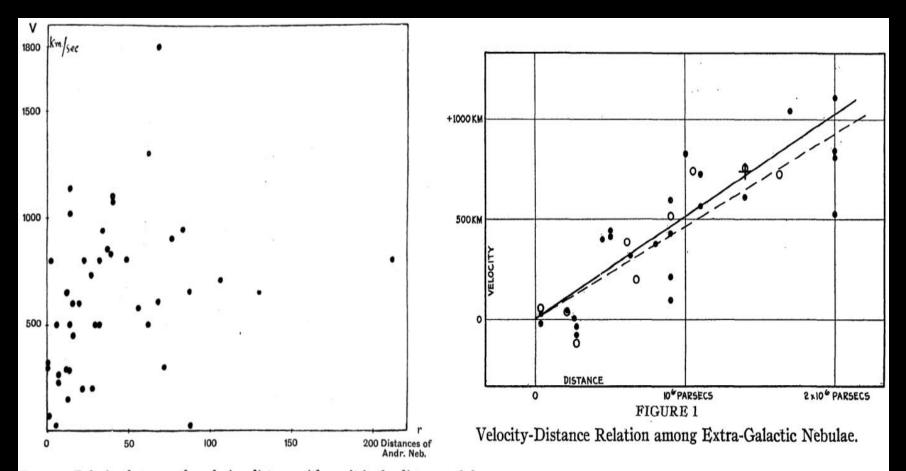


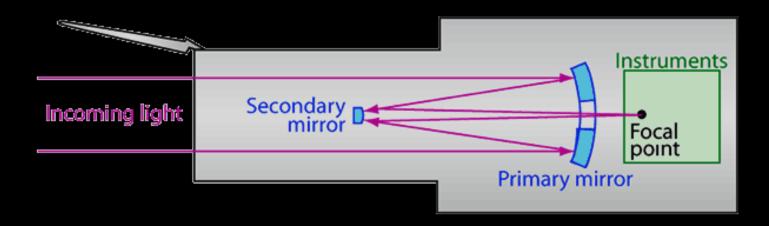
Fig. 5.—Relation between the relative distances (the unit is the distance of the Andromeda nebula) and the measured radial velocities of spiral nebulæ.

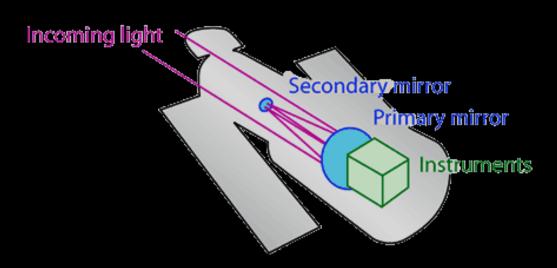


Hubble

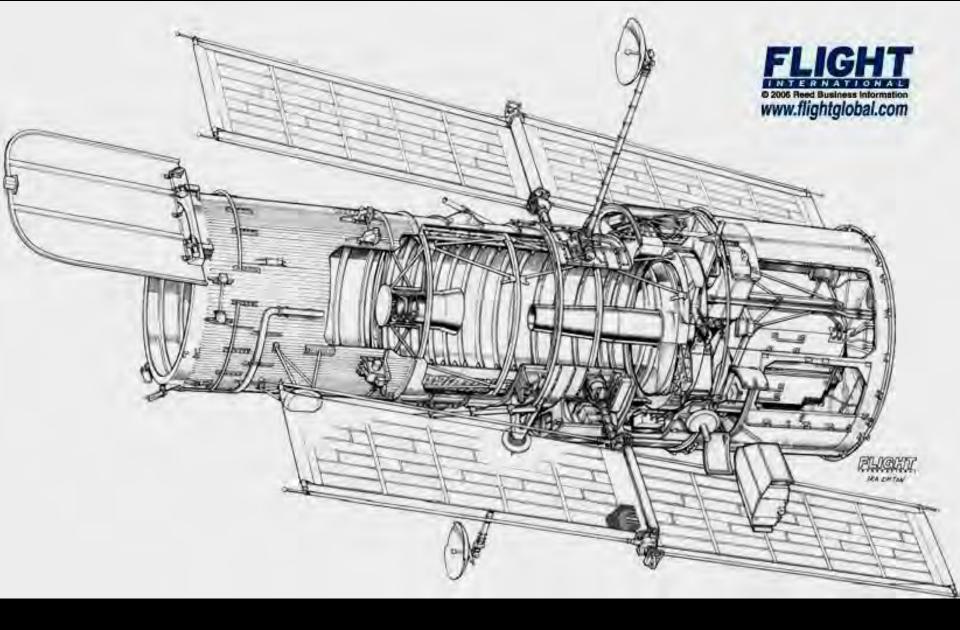
the Telescope

NASA/ESA project, to provide an optical high resolution observatory in LEO

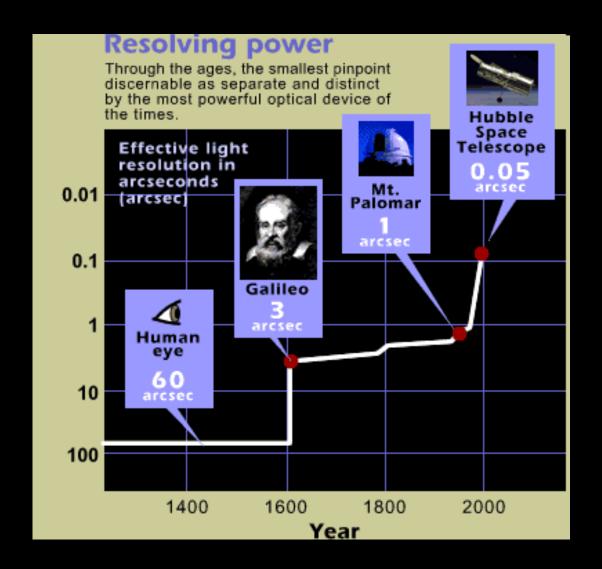




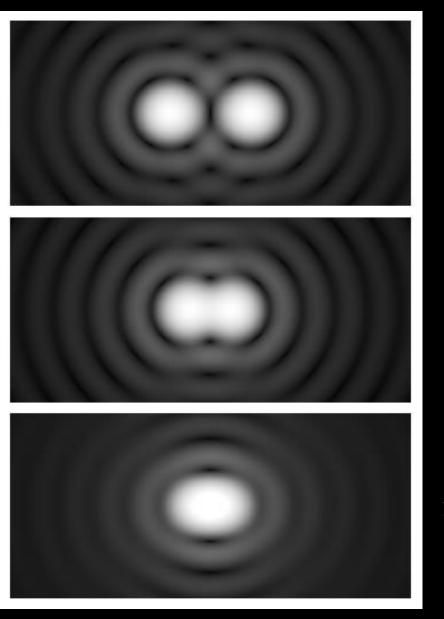
Hubble optical system



The Telescope



Expected resolution of the Telescope, compared to other optical systems



resolved

just resolved

not resolved

Airy diffraction patterns generated by light from two points passing through a circular aperture

The angular resolution R of a telescope can be approximated by

 $\mathbf{R} = \lambda/\mathbf{D}$

 λ is the wavelength of the observed radiation and D is the diameter of the telescope's objective or mirror.

Resulting R is in radians. Sources larger than the angular resolution are called extended sources or diffuse sources, and smaller sources are called point sources (stars are obviously point sources).

For example, in the case of yellow light with a wavelength of 580 nm, for a resolution of 0.1 arc second, we need D = 1.2 m.

For Hubble, we have D = 2.4 m, so that the resolution is





A very carefully polished primary mirror, but there was a problem, not discovered until on-orbit...





Shuttle Cockpit





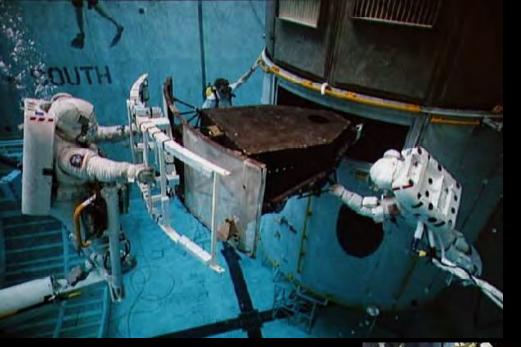
Servicing Missions (SM) from 1993 to 2009



Crew of HST SM 1 (STS-61), 1993

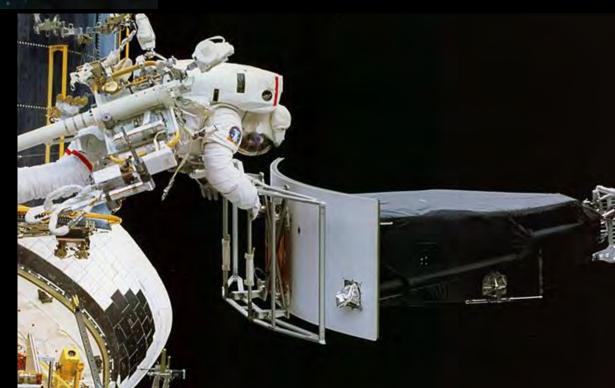


Evaluation of servicing strategies for SM 1



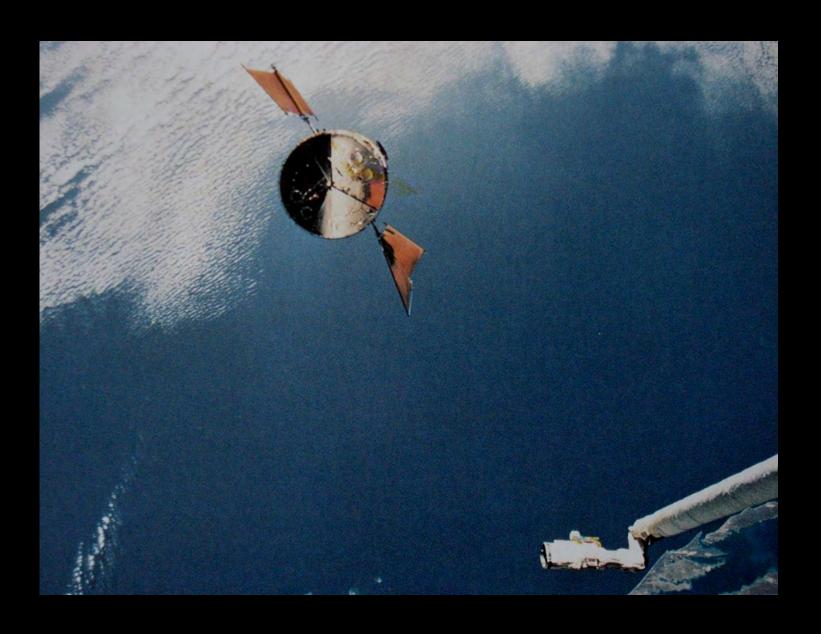
SM 1 training in the water...
MSFC Huntsville, Al.

...and reality



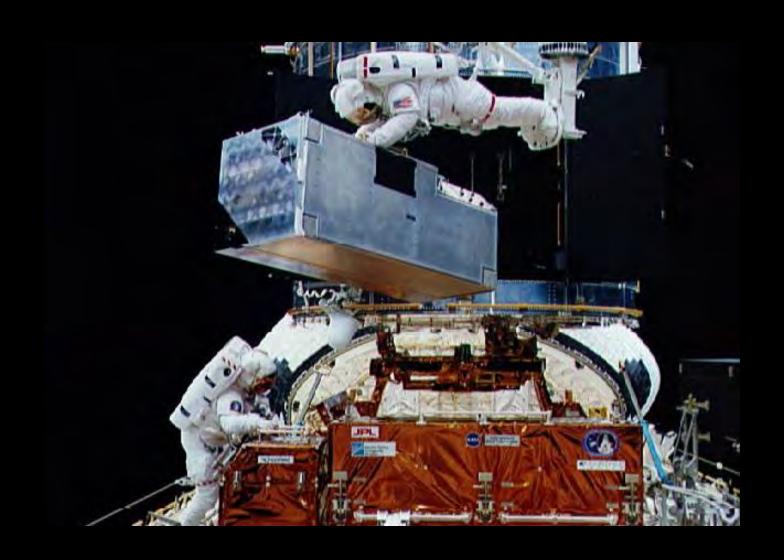


Lift-off - December 2, 1993 - 4:26AM

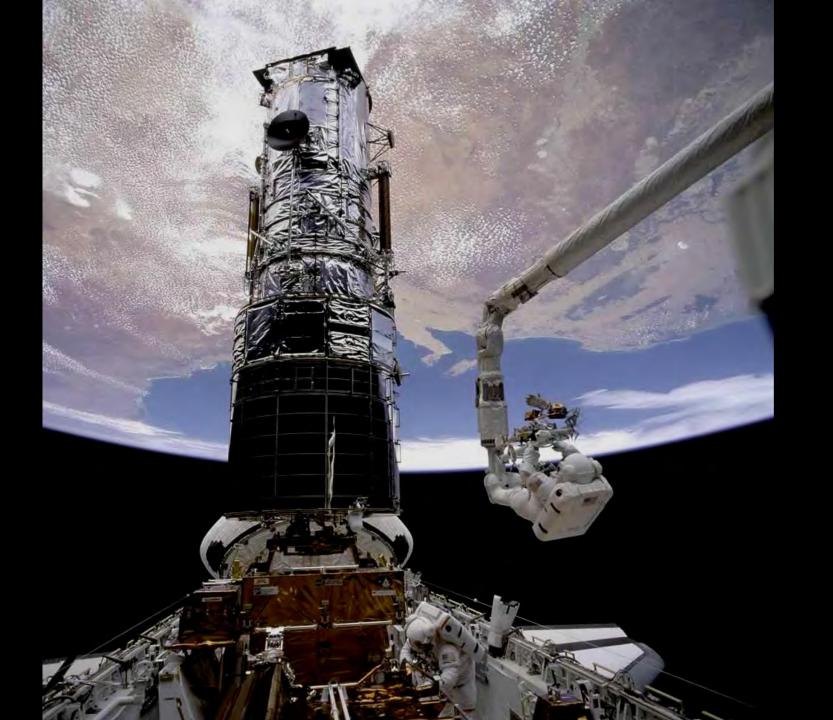


We found it!





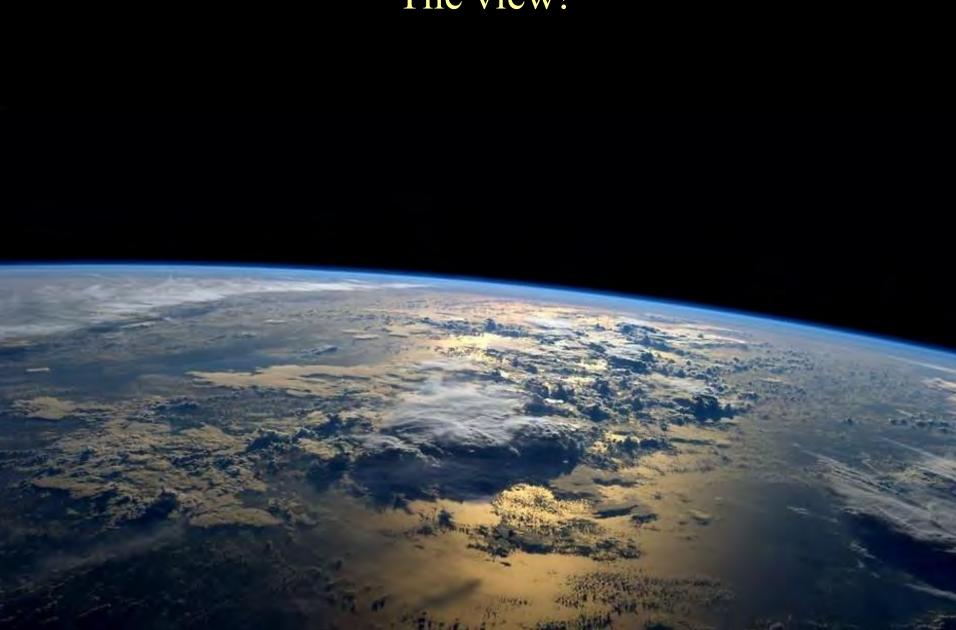
COSTAR installation





A happy crew!

The view!









Reentry... HOT!





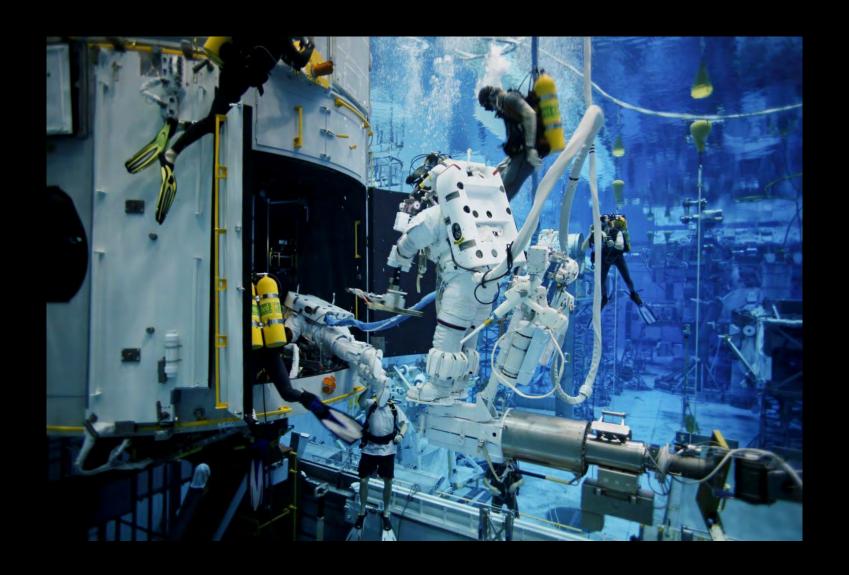


Milt Heflin



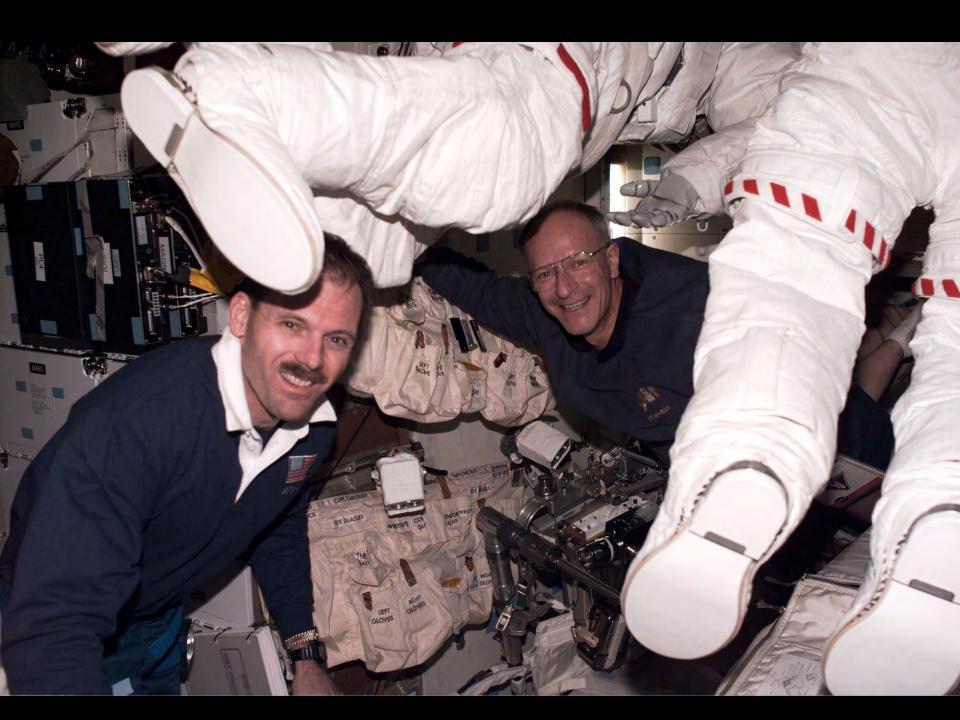
Lead Flight Director

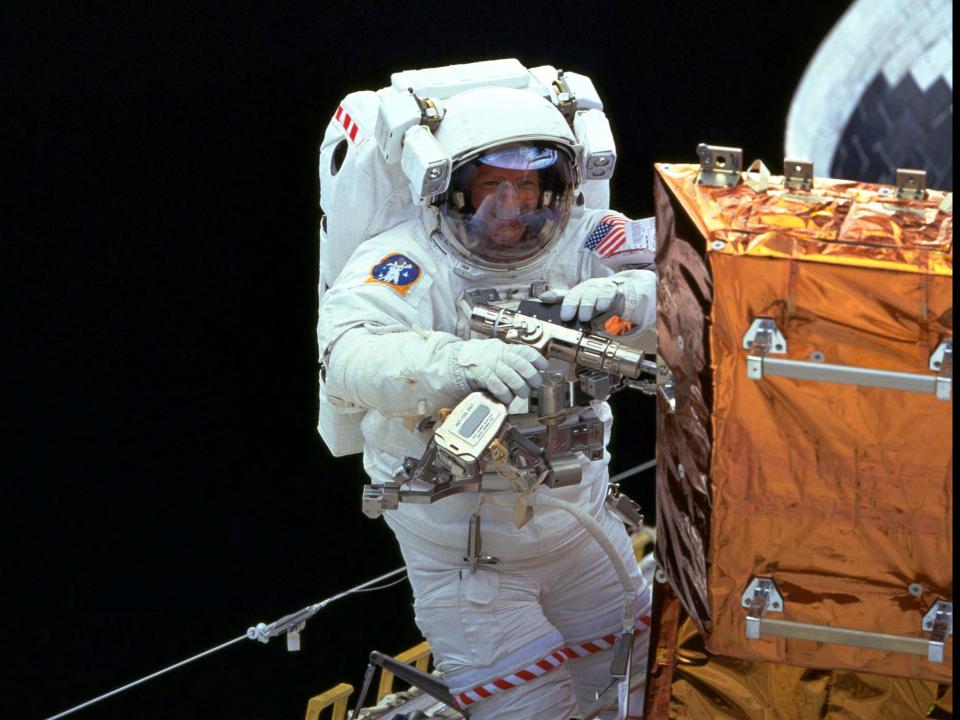
Hubble First Servicing Mission



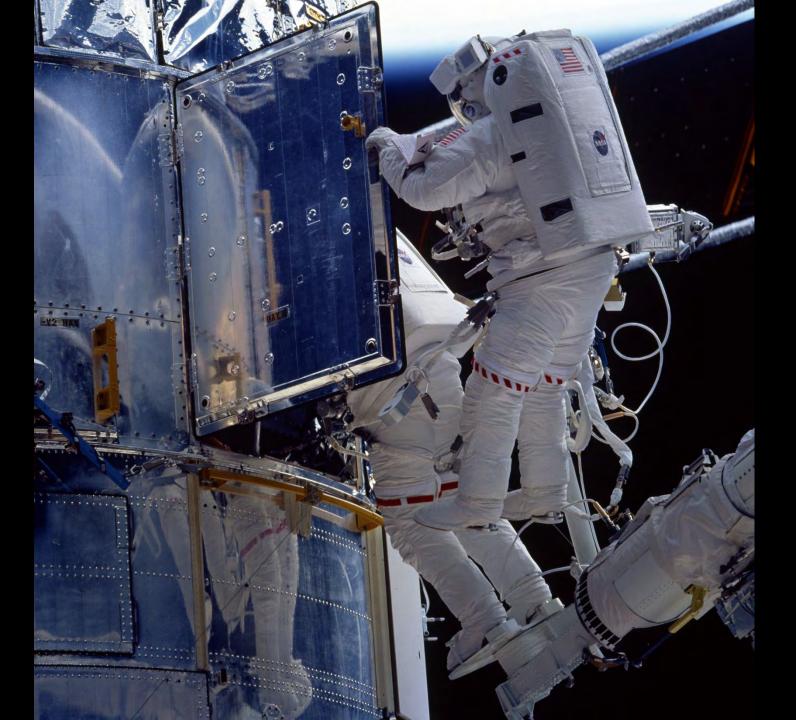
SM 3A Training



















...it was Christmas Day, 1999!



Hubble SM 4 crew







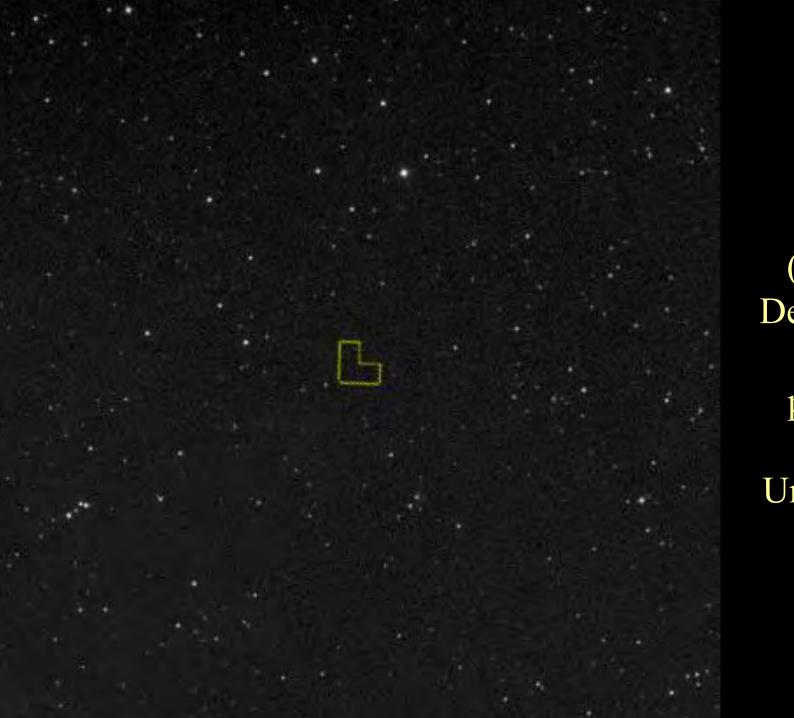
Cat's Eye Nebula • NGC 6543





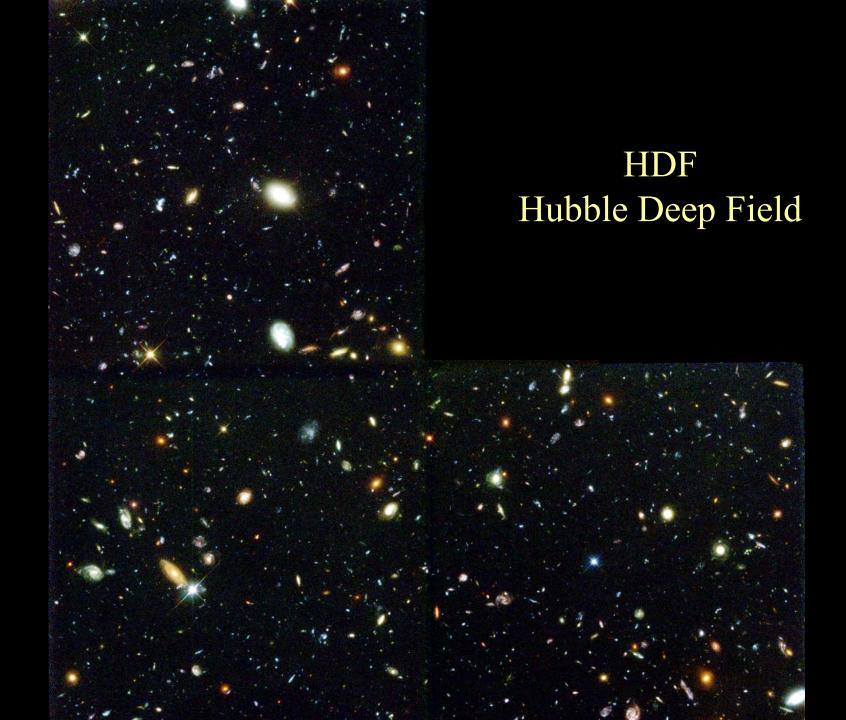
Interacting Galaxies Arp 273





HDF (Hubble Deep Field)

position in Ursa Major



Why so much success in HST servicing?

- Clear goals and priorities
- Strict operational discipline, in space and on the ground
- Teamwork
- Thorough training in realistic environments extensive practice for handling malfunctions

James Webb Space Telescope (JWST)



Until (and hopefully beyond) JWST, we hope that we will enjoy many more discoveries and stunning pictures from the Hubble Space Telescope!

Thank you for your attention!