Space Science A Tool for International Cooperation

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International Space Science Institute
Bern



Main points

- What is space science?
- The main players
- Is International cooperation necessary?
- How did International Cooperation start?
- How to cooperate?
- How not to cooperate?
- The future of International Cooperation



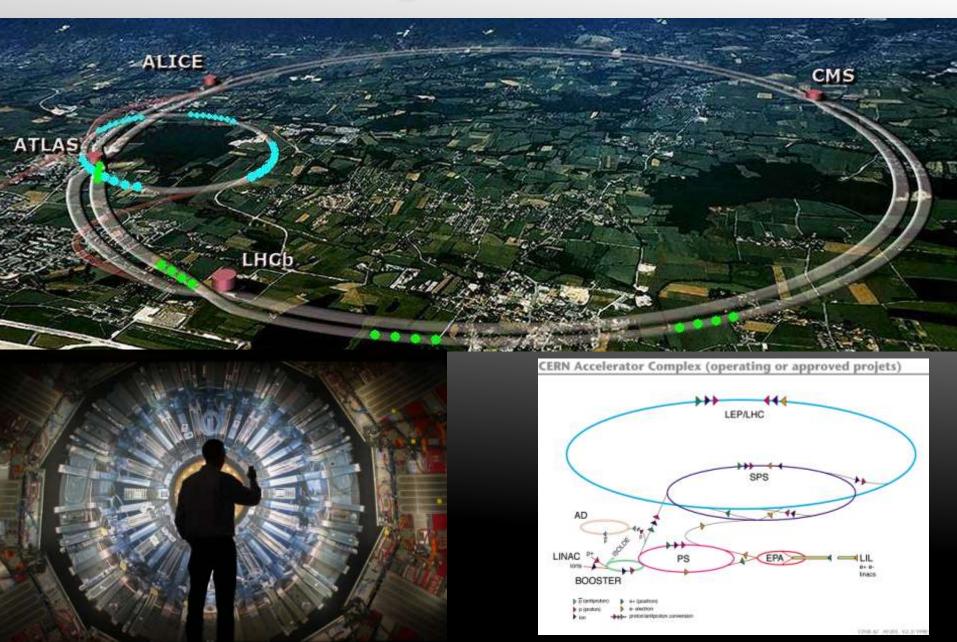
What is Space Science?

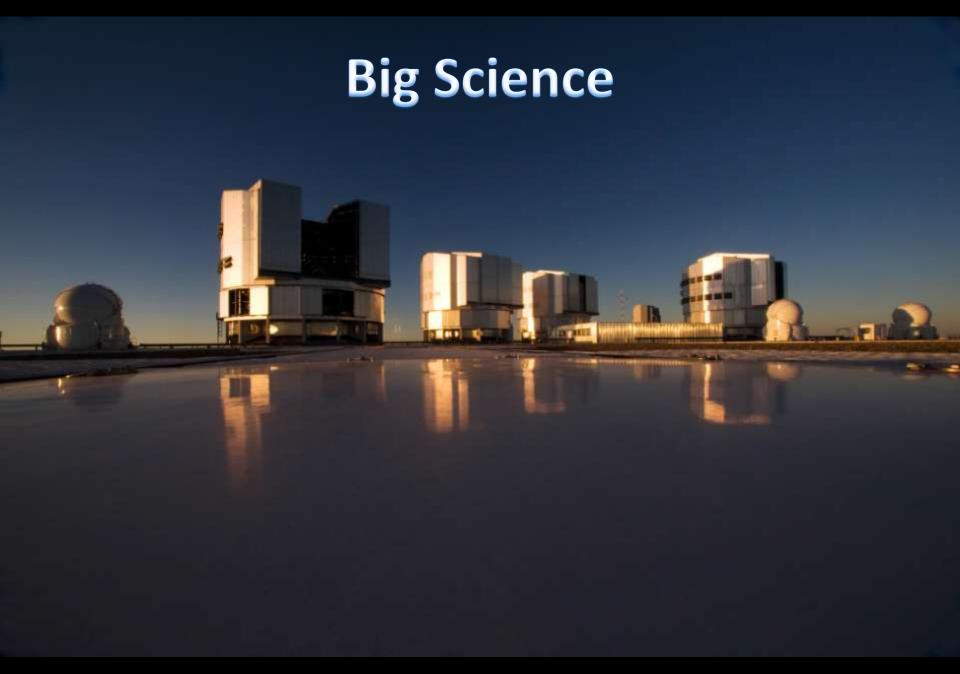
- Astronomy and fundamental physics
- Solar System exploration
- Plasma physics
- Earth sciences

Space Science is "Big Science"



Big Science







Big Science

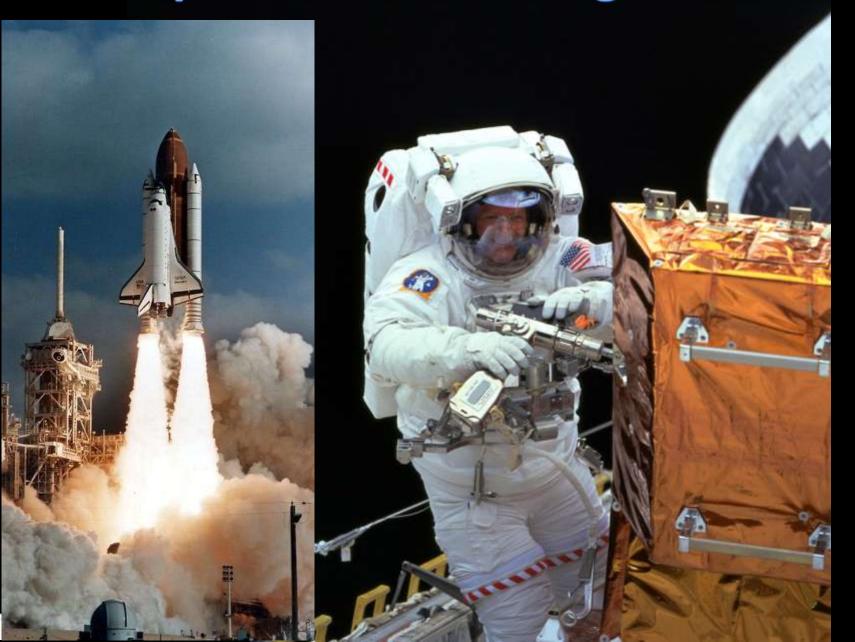


Big Science Tianyan (天眼)





Space Science is Big Science



The main space organizations

- APSCO
- China NSSC, CAS NSSC
- India ISRO इस्ती
- Japan JAXA







Thaïland NSPO





United States NASA Mag





Is there any space science endeavor without IC?

- "Autonomous science" is generally not the best science!
- Science works only best when confronted to the broadest basis of knowledge!
- Science (space science) has always been, is, and will continue to be INTERNATIONAL!
- Hence, best science is dependent upon the broadest international confrontation
- Independence is a political approach, not a science approach!



What projects are concerned

Mostly B\$ projects



 Small projects also but they might be better fit to independent endeavors





Why cooperating?

- Sharing the results of science discoveries
- Some missions may be too big or too costly for one organization alone to build
- Necessary talents may not exist in the principal country
- Bringing added value and complementary expertise



Sharing the results scientific discoveries

Dust measurements in the coma of comet 67P/Churyumov-Gerasimenko inbound to the Sun

Alessandra Rotundi, 1,80 Holger Sierka, Vincenzo Della Corte, 1 Marco Fulle, 4 Pedro J. Gutierrez, Luisa Lara, Cesare Barbieri, Philippe L. Lamy, Rafael Rodrigo, 8.8 Detlef Koschny, 10 Hans Rickman, 11,13 Horst Uwe Keller, 13 José J. López-Moreno, 2 Mario Accolla, 1,2 Jessica Agarwal, 2 Michael F. A'Hearn, 14 Nicolas Altohelli, 15 Francesco Augrilli, 16 M. Antonietta Barucci, 17 Jean-Loup Bertaux, 18 Ivano Bertini, 16 Dennis Bodewits,34 Erio Bussoletti,3 Luigi Colangeli,36 Massimo Cosi,31 Gabriele Cremonese,22 Jean-Francois Crifo,18 Vania Da Deppo,25 Björn Davidsson,11 Stefano Debei, 16 Mariolino De Cocco, 26 Francesca Esposito, 26 Marco Ferrari, 1,2 Sonia Fornazier,17 Frank Giovane,26 Bo Gustafson,27 Simon F. Green,26 Olivier Groussin,7 Eberbard Grün, 26 Carsten Güttler, Migsel L. Herranz, Stubbe F. Hviid, 30 Wing Ip, 31 Stavro Ivanovski, 1 José M. Jerónimo, 1 Laurent Jorda, 1 Joerg Knollenberg, 50 Rainer Kramm,5 Ekkehard Kührt,20 Michael Küppers,15 Monica Lazzurin,6 Mark R. Leese, 26 Antonio C. López-Jiménez, Francesca Lucarelli, 2 Stephen C. Lowry, 32 Francesco Marzari, AN Elena Mazzotta Epifani, N. J. Anthony M. McDonnell, Sa., 13, 34 Vito Mennella, 28 Harald Michalik, 25 Antonio Molina, 26 Rafael Morales, 5 Fernando Moreno, Stefano Mottola, 30 Giampiero Naletto, 37 Nilda Oklay, 3 José L. Ortiz, 8 Ernesto Palumba, 1 Pasquale Palumbo, 1,4 Jean-Marie Perrin, 18,38 Julio Rodriguez, 5 Lola Saban, 20 Colin Snodgrass, 5,20 Roberto Sordini, 1 Nicolas Thomas, 40 Cecilia Tubiana, Jean-Baptiste Vincent, Paul Weissman, Klaus-Peter Wenzel, III Vladimir Zakharov, 17 John C. Zarnecki 6,38

Critical measurements for understanding accretion and the dust/gas ratio in the solar nebula, where planets were forming 4.5 billion years ago, are being obtained by the GIADA (Grain impact Analyser and Dust Accumulator) experiment on the European Space Agency's Rosetta spececraft orbiting comet 67P/Churyumov-Gerasimenko. Between 3.6 and 3.4 astronomical units inbound, GIADA and OSIRIS (Optical, Spectroscopic, and Infrared Remote Imaging System) detected 35 outflowing grains of mass 10⁻¹⁰ to 10⁻⁷ kilograms, and 48 grains of mass 10⁻⁵ to 10⁻² kilograms, respectively. Combined with gas data from the MIRO (Microwaye Instrument for the Rosetta Orbiter) and ROSINA (Rosetta Orbiter Spectrometer for ion and Neutral Analysis) instruments, we find a dust/ gas mass ratio of 4 * 2 averaged over the sunlit nucleus surface. A cloud of larger grains also encircles the nucleus in bound orbits from the previous perihelion. The largest orbiting clumps are meter-sized, confirming the dust/gas ratio of 3 inferred at perihelion from models of dust comae and trails.

Ithough accurate measurements of the gas loss rate from comets are possible under favorable conditions even from Earth, estimates of the dust loss rate have been more uncertain thus far. Multiparametric models are needed to extract global dust parameters from the dust features of comets (coma, tail, and trail) observed from the ground and Earthorbiting telescopes, and it is often difficult to establish the uniqueness of these model results. Past space missions had onboard instruments devoted to the measurement of the dust flux. Because all of these missions were fast flybys, it was impossible to disentangle the dust grains coming directly from the nucleus from those reflected back by solar radiation pressure (1, 2). The latter component could explain part of the excess of millimeter-sized particles, showing as a bulge of the size distribution above the fitting power law derived by the observations performed during flybys at IP/Halley (3) and at short-period

comets 26P/Grigg-Skjellerup, 81P/Wild 2, and 9P/Tempel 1 (#-6).

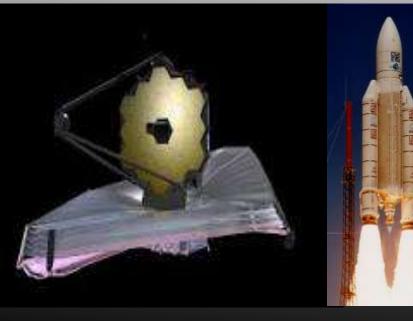
An even more severe bias could affect all estimates of dust/gas ratio obtained so far in comets. The dust/gas ratio measured in 1P/Halley was close to 2 (5), but this number is valid up to the largest mass of ~1 g observed by the DIDSY (Dust Impact Detection System) detector (actually, this largest-mass grain was invoked to explain the spacecraft precession-inducing impact that occurred just before closest approach). Theoretical models predict that 1P/Halley was then ejecting larger masses, and the dost/gas ratio strongly depends on the actual largest grain ejected in the coma. Because it was impossible to fix the size distribution between 1 g and the unknown largest ejected mass, we cannot exclude dust/gas values a factor of 10 higher. In this paper, we show that for comet 67P/Churyumov-Geneimmko (67P hereafter), we can disentangle the two families of ejected grains (direct and reflected) and extract the dust size distribution up to the largest ejected grain, obtaining for the first time an accurate estimate of the dust/gas ratio.

Istituto di Astrofisica e Planetologia Scaniali, Istituto Nazionale di Astrofisica (INAF), Via Fosso del Cavaliere, 200. 0035 Rome, Baty, ²Groverstá degli Stydi di Nopoli "Partherape," Dipartimento di Scienze e Tecnologie, CON IC4, 80343 Naples, Italy. "Max Planck Institut für Somereyatenforactung, Austra-von Lieby Weg, 3, 37077 Gottingen, Germany, "Osservaturio Astronomico, BNAF, Visa Tiepolo II, 34043 Eneste, Italy. Trabitato de Astrofísica de Andriucia, Cerusjo Superior de Investigaciones Certificas. (CSIC), P.O. Box 2006, 35080 Granada, Spain, *Department of Physics and Astronomy, Padava University, Vicsili dell'Opporvatorio S. 25022 Padinia, Nolv. Laboratore ∉'Autrophysique de Marsette, Utdit 7326. CNRS and Air-Marcelle Orwanski. 13365 Marcelle, France. *Cantro de Autrobiologia (instituto Nacional de Tecnica Aerospocial-CSIC), 29601 Villanueue de la Canada, Madrid, Scoin, *International Space Science Institute, Hallendroom 6, CH 2012 flors, Switzerland, ³⁰Scientific Support Office, European Space Agency, 2201 Noordwijk, Notherlands, ¹³Department of Physics and Astronomy, Goossis University, Box 516, 75120 Sponses, Sweden, "Polish Academy of Sciences Space Research Certier, Bartycka IBA, Pt.-00716 Warszaws, Poland. Snotitute for Geographics and Extratementrial Physics, Sectimate Universität Braunstiweig, Braunstiweig 38106, Germany: PiDepartment of Astronomy, University of Maryland: College Park, MD 30742-3421, USA. "European Space Agency (ESA)-European Space Astronomy Center (ESAC), Camino Bigo del Castillo, s/n. 28692 Villanueve de to Cahota, Moorld, Spagno, ¹⁶Cirpartment of Mechanical Engineering, Liniversity of Pagena, via Venezia 1, 261(3) Padova, Raty. "Laboratore d'Etudes Spatiales et. d'Instrumentation en Astrophysique, Observatoire de Paris. CNRS, Université Paris, 06, Université Paris-Cidenti, 5 place Johannes Janssen, 92195 Meudon, France. "Laberatoire: Atmosphines, Mileux, Observations Spotistes, CNRS/ Université de Versailles Saint-Quentin-en-Yvelines/fratiful Purre-Conon Laptice. IJ boulevard of Riembert, 790390. Guyanosurt, France. "University of Pastova, Centro Interdipartimentale di Studi e Alfarità Spazzali (CESAS), via Venezia 15, 35100 Padova, Italy, 19ESA, European Space Research and Technology Centre (ESTEC), Keptertain 1. 2201 AZ Noordwija, Netherlands, ⁵¹Sales ES, Via Albert Einstein, 35, 50013 Corrue Bismois, Firence, Haly. Osseniatorio Astronomico di Padova, INAF, Vicalia didf Disservator is 5, 35622 Partiess, Maly: 25 Directation Niglorate delle Ricerche-latituta di Fatorica e Nonotecrologie-Unità Operativa di Supporto Passiva LURDR. via Tromes 7, 20131 Pactous, Raly. "Università di Trento, via Mesiano, 77, 38100 Trento, Raty, 20 Osservatorio Astronomico si Capodimortis, INAF, Selto Molariello, 16, 80133 Naples. Buly. **Wigers Polytechnic Institute and State University. Standary, VA 24003, USA. **University of Florida. Generatio, Ft. 3260, USA. Styllawday and Space Sciences. Department of Physical Sciences, The Open University, Million Keynes 5817 SAA, UK. ²⁵Mox-Planck-institut fuer-Karrystynik, Sauphercheckung 1, 69017 Heithelberg, Germany. Minstitute of Planetary Research, Deutsches Ziertrum für Early and Reventerst (DLR), Rultowforcetrome 2, 12489 Bierkin, Germany. Mitrotifule for Space Science, National. Central University, 300 Chang Da Road, 32054 Chang Li. Taiwan. 28 The University of Kent, School of Physical Sciences, Cartectury, Kent CT2 7NZ, CK. SSppartment of Physics, University of Padova, 25131 Padova, Italy. ³⁶Grespacekent, Centerbury C22 BEF, UK. ³⁶Institut für Datemachnik und Kommunikationsreitze, 38006 Braunschweig, Germany, MOepartomento de Fisica Agricada. Universidad de Granada, Facultad de Ciencias, Avenida Severo Ochsa, s./n, 18071 Granada, Spain. "Department of Information Engineering, Padova University, via Graderigo 6. 35131 Padovo, Raly. 35Observatoire de Haute Presence OSU Pythess UMS 2244 CNRS-AMU, 04676 Saint Michel l'Observatoire, France, ²⁸institute Nacional de Teorica. Aeroenpasial, Carretera de Ajakir, p.h. 4, 25850 Torrejon de Arctor, Madrid, Spain. 40 Physikalisches Institut, Scientrasse 5. University of Dem, 3002 Bern, Switzertand, 4 Planetary Science Section, Jet Propulsion Laboratory, 4600 Oak Grove. Drive, Passateria, CA 95109, USA. *Corresponding author: E-mail: rotundiffuniportherape it



Some missions may be too big or too costly for one organization alone to build





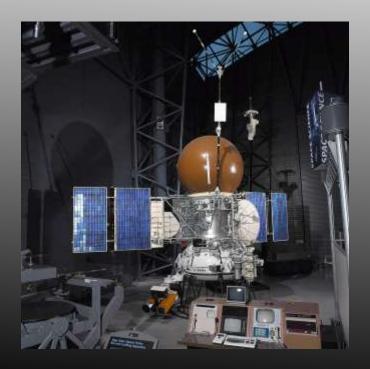
JWST

ISS

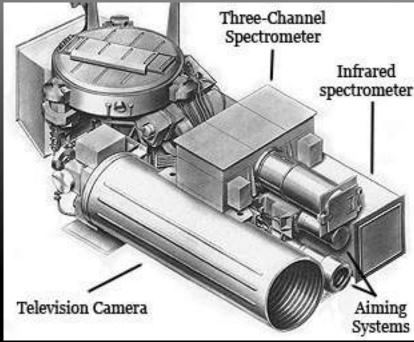


Necessary talents may not exist in the principal country

VEGA

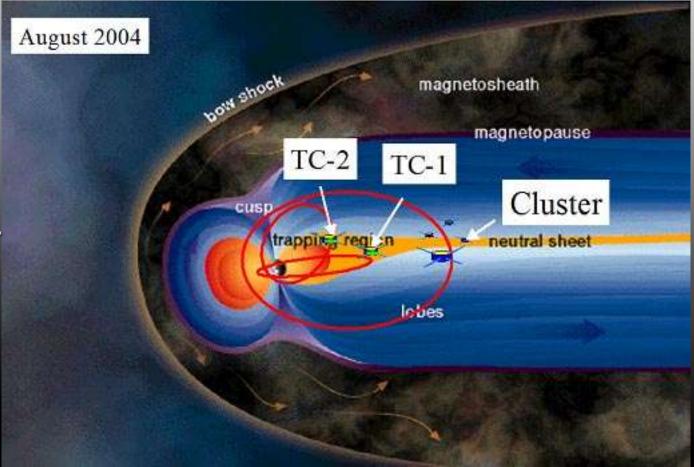


IKS



Bringing added value and complementary expertise

- DAMPE
- Double Star
- QUANTUM





IAA Team Achievement Award 2010



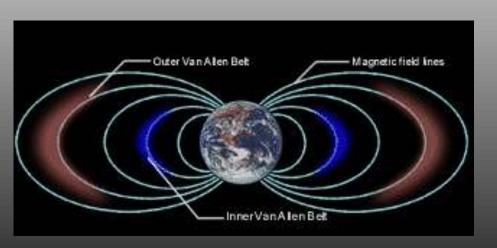


How did it start?

- In the difficult context of the Cold War after the launch of Sputnik 1
- Competition was strong between the two super powers













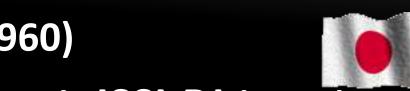
How did it start?

Soon the space and scientific community created means of cooperation:



(Beijing-2006, Mysore-2012, Leningrad-1970,Tokyo-1968, Nagoya, 1998

• IAA (1960)









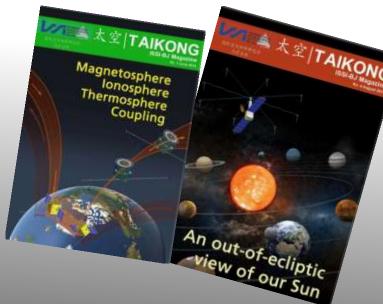












- International Teams
- Workshops
- Working groups
- Forums
- Summer Schools
- Visiting scientists



Space Very Long Baseline Interferometry Forum



How to cooperate?



How to cooperate?

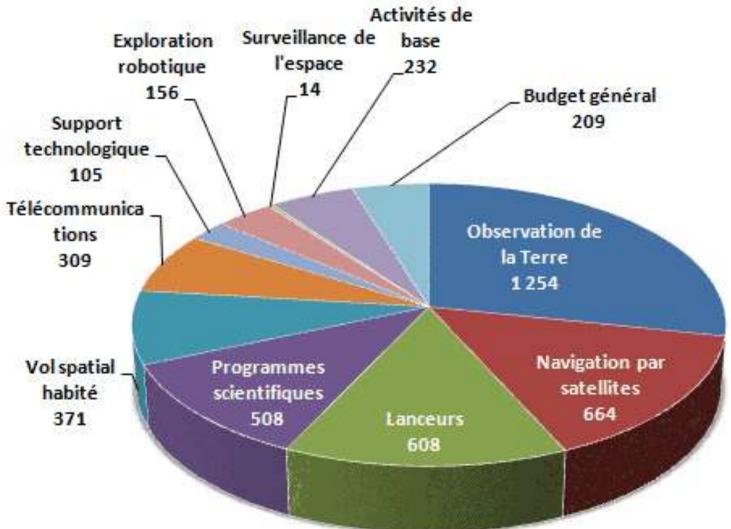
- The ESA approach
- Mutual interest through complementary contributions from the partners
- One leading partner is usually needed
- Mutual respect among the partners is a mandatory rule for a successful result
- Examples of various cooperative missions







ESA Programs and respective Budgets (Billions €)

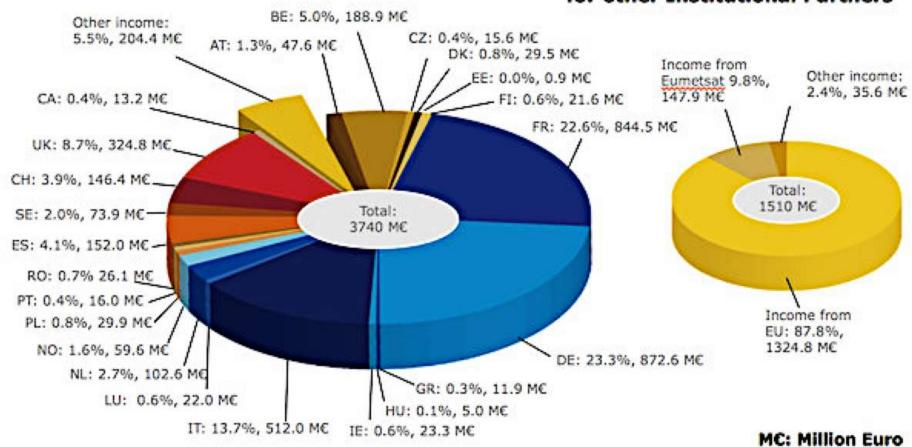




ESA Budget for 2016

ESA Activities and Programmes

Programmes implemented for other Institutional Partners



TOTAL ESA BUDGET FOR 2016: 5250 M€

European Space Agency

The ESA approach to International cooperation

- Concerns mostly missions too big for a single European nation to develop
- Share of responsibility between Member states and ESA:
 - ESA provides the common systems (Spacecraft, Launcher, in orbit operations, data delivery,
 - Member states fund all expenditures related to these parts of the scientific payload developed in their respective country
- "Coopetition": Member states ESA contributions are in competition with their domestic activities







Hardware contribution HST





Hardware contribution SoHO

- Joint program between the European Space Agency (ESA) and NASA.
 - An industry team led by Matra Marconi Space built SOHO in Europe. It's instruments were provided by nine European and three U.S. Principle Investigators
 - ESA: responsible for SOHO's procurement, integration, and testing
 - NASA: provided launch and mission operations (at Goddard Space Flight Center)

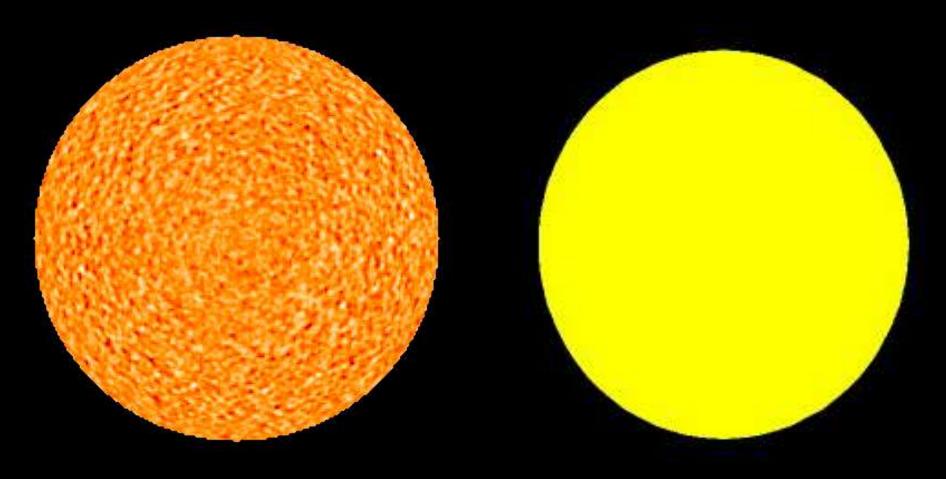




Launch 2 December 1995

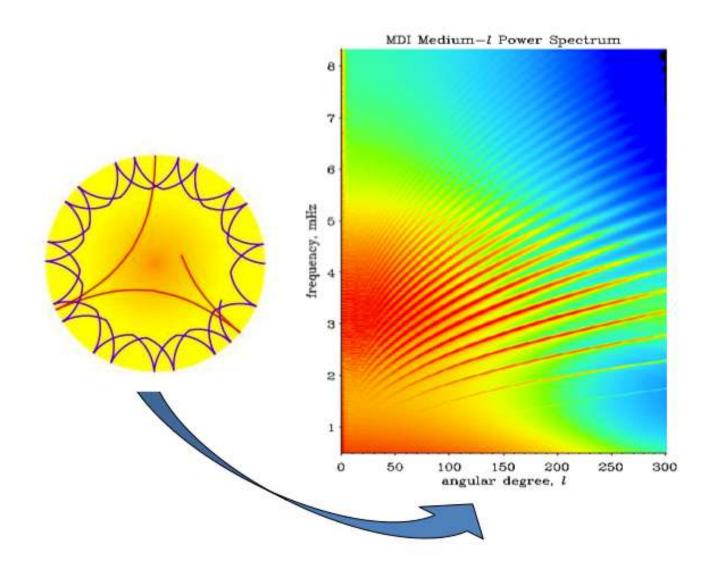


Probing the Sun's interior





Probing the Sun's interior



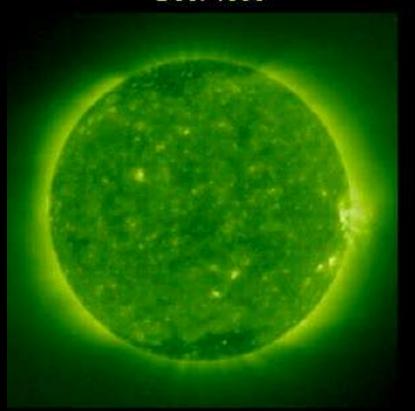


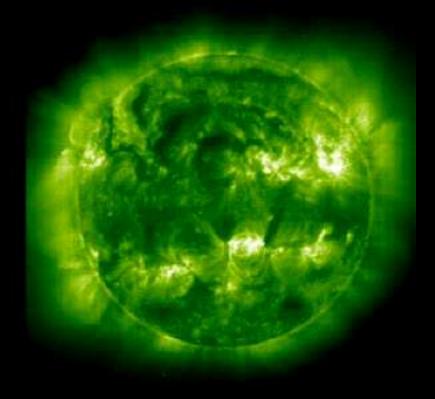
Comets and CMEs



The solar corona as seen from SOHO

EIT 195 Å Dec. 1996 EIT 195 Å June 1999







Complementarity

XMM-Newton ESA

Chandra- NASA



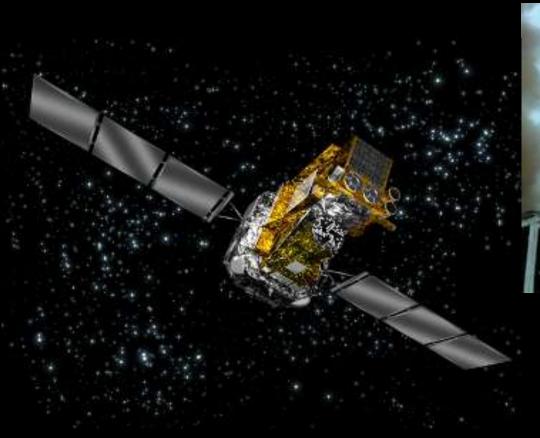








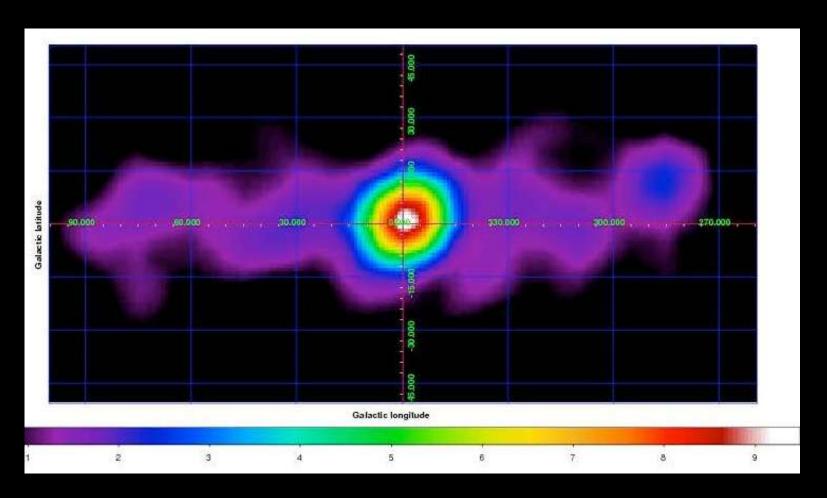
Multilateral agreements INTEGRAL (ESA-Russia-NASA) INTErnational Gamma-Ray Astrophysics Laboratory







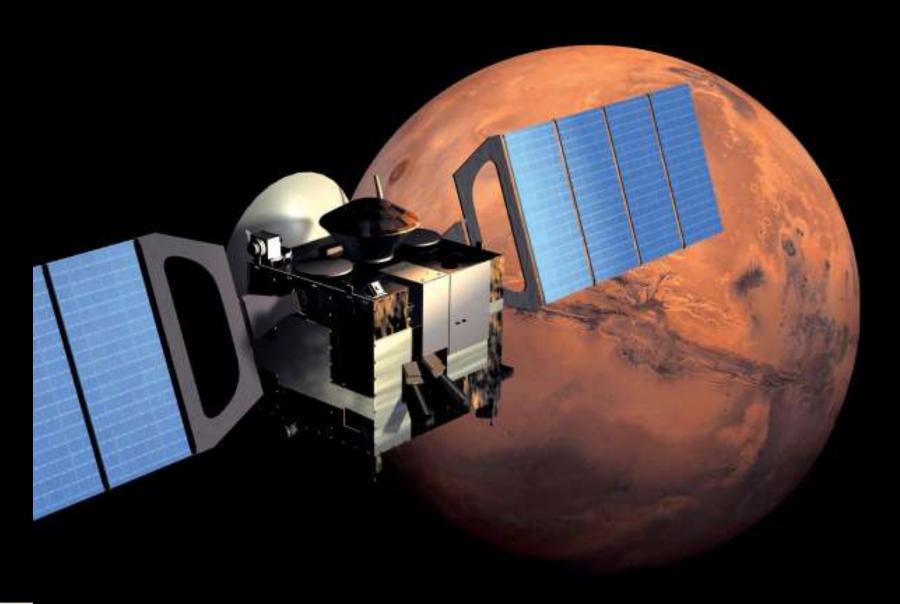
INTEGRAL



Spatial distribution of the 511 KeV positron annihilation emission in the galactic center

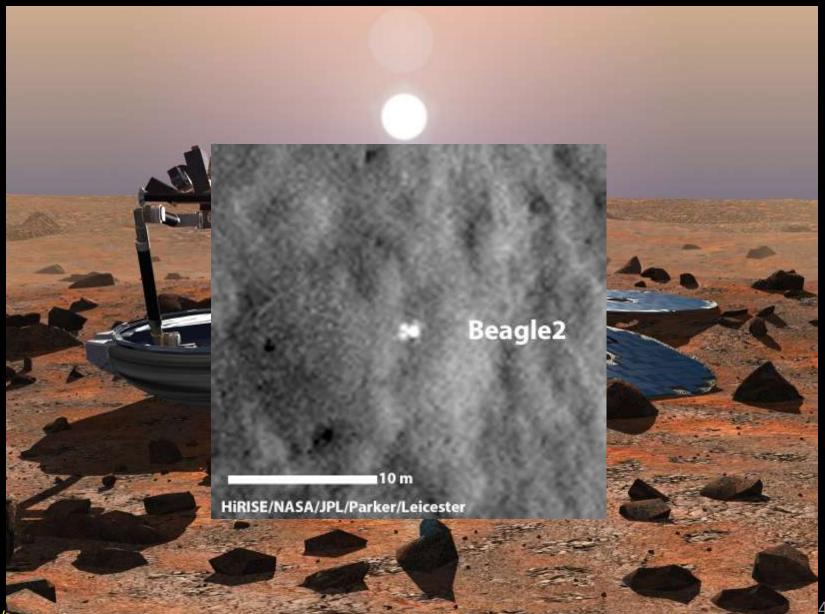


Mars Express and Beagle 2





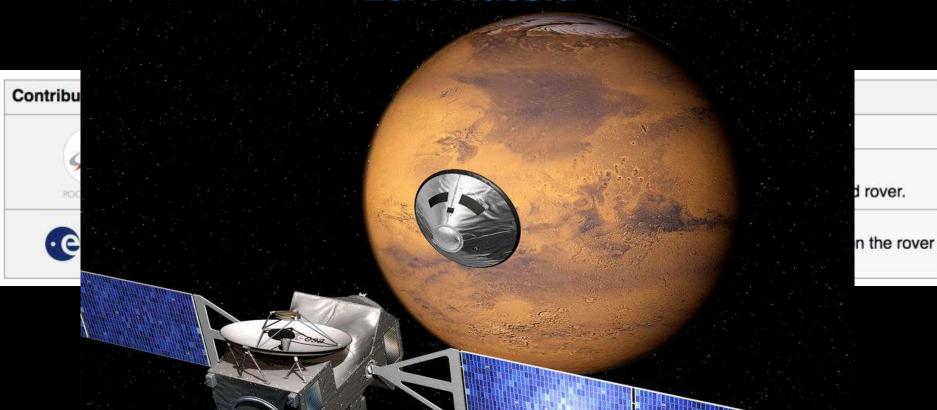
Mars Express and Beagle 2





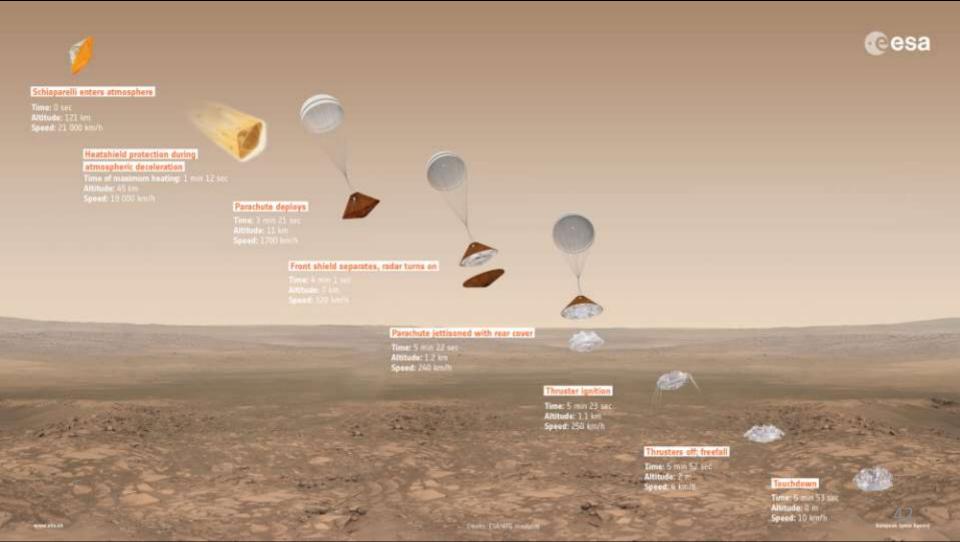
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ExoMars 3кзоМарс ESA-Russia



41

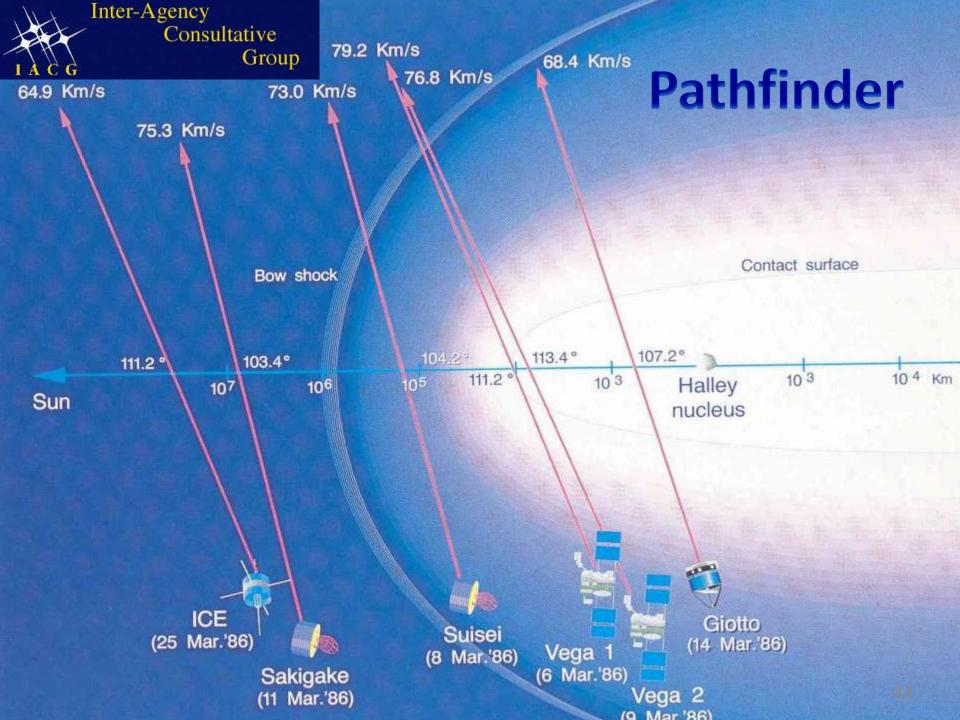
ExoMars 3кзоМарс ESA-Russia



Cooperation through coordination

- Giotto-Halley pathfinder concept
- ILWS
- Earth sciences









International Living With a Star

SCIENCE

CHARTER

ORGANIZATION

MISSIONS

NEWS

PRESENTATIONS

HISTORY

CONTACTS







ILWS Charter

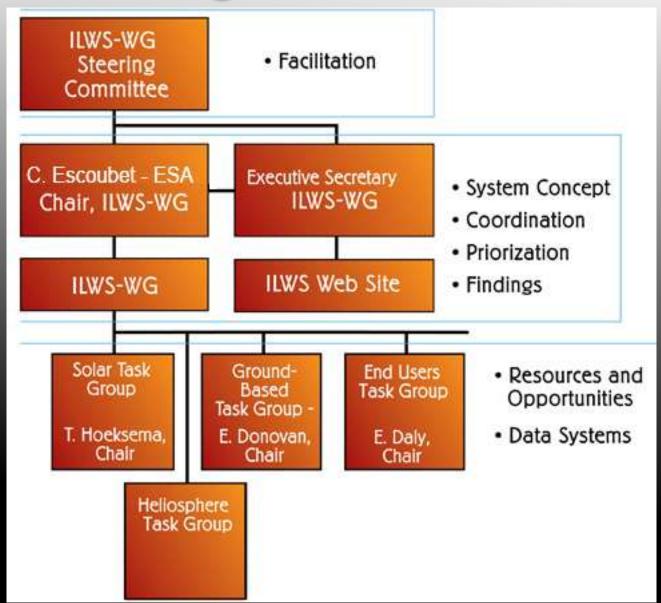
- Stimulate, strengthen, and coordinate space research to understand the governing processes of the connected Sun-Earth system as an integrated entity.
- Objectives
 To stimulate and facilitate:
- Study of the Sun-Earth connected system as a system and its effect which influence life and society.
- Synergistic coordination of international research in solarterrestrial physics, including all relevant data sources as well as theory and modeling.
- Collaboration among potential partners in solar-terrestrial space missions.
- Effective and user driven access to all data, results, and valueadded products.



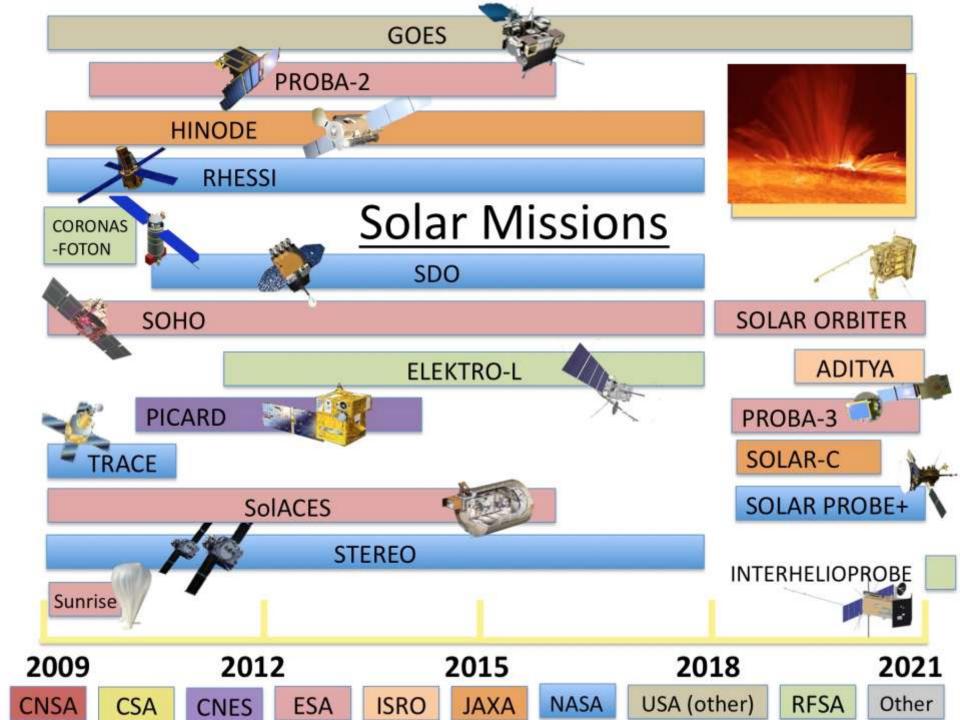


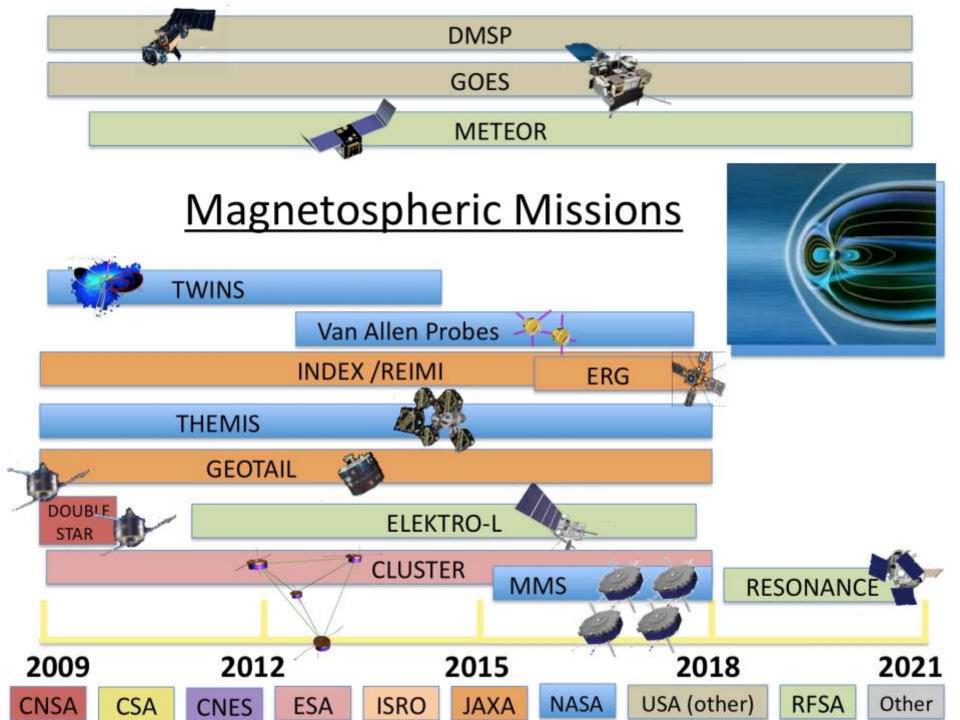


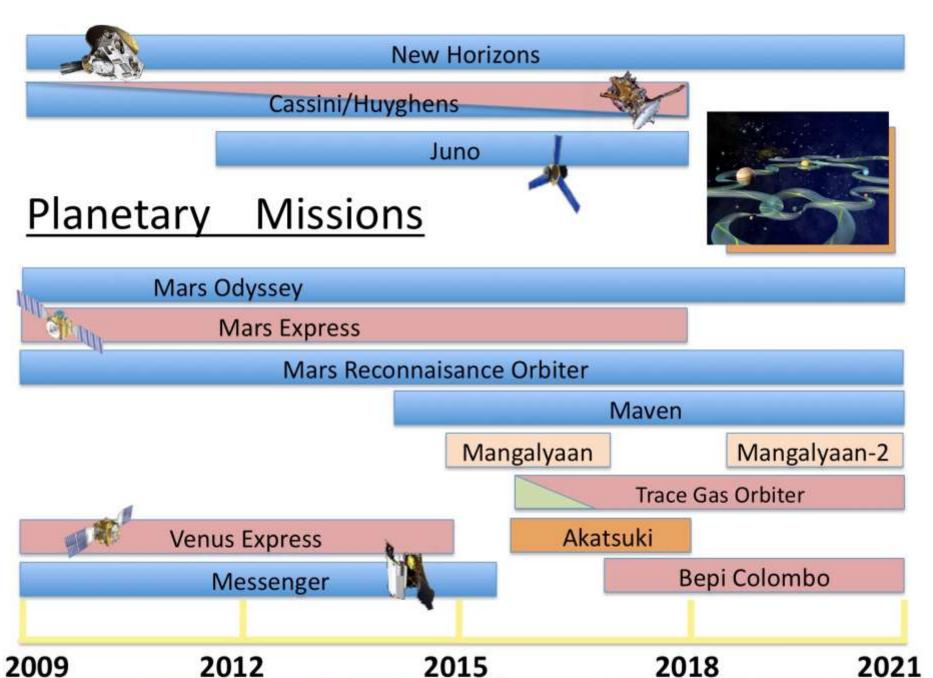
Organization











CNSA CSA CNES ESA ISRO JAXA NASA USA (other) RFSA Other

The special case of Earth observation satellites

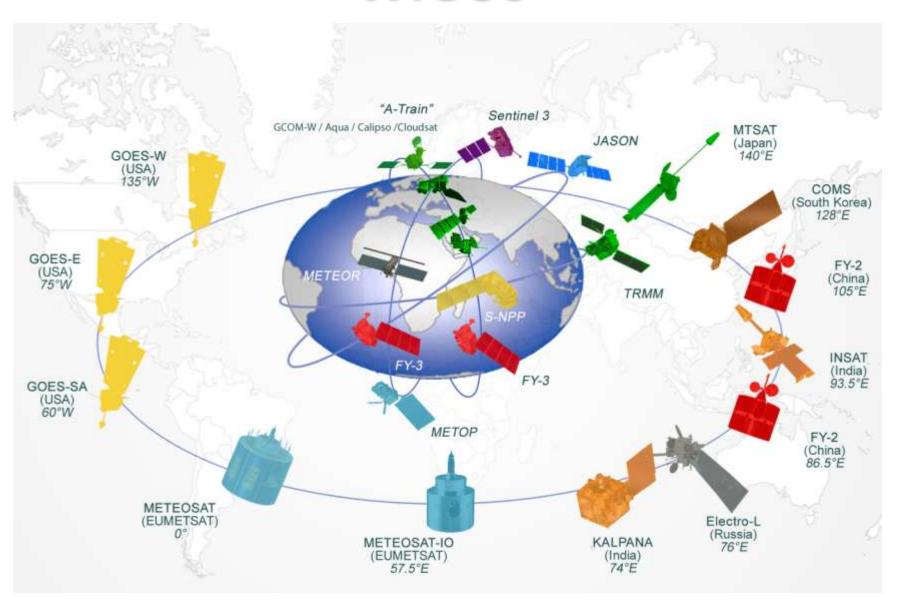


CEOS Background

- Optimize the benefits of spaceborne Earth observations through cooperation in mission planning and development of compatible data products, formats, services, applications & policies;
- To serve as a focal point for international coordination of space-related Earth observation activities;
- To exchange policy and technical information to encourage complementarity and compatibility among spaceborne Earth observation systems and the data received from them.



WIGOS

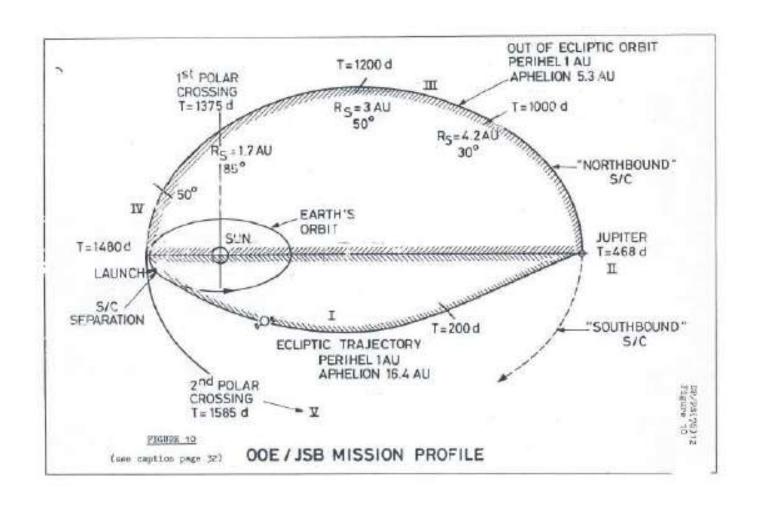




How not to cooperate



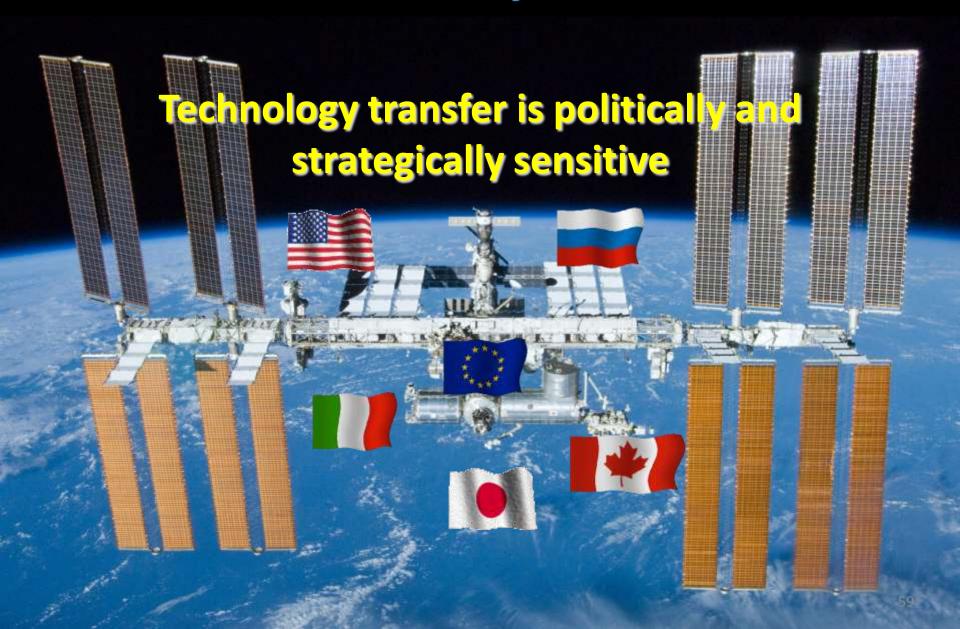
ISPM Original concept







International Space Station



Conclusions



Some general rules

- International cooperation should be the tool for the development of an optimum worldwide program for the benefit of space science.
- Programs of the different agencies should be complementary to one another and not compete, while allowing a healthy overlapping.
- Agencies should exercise fairness, restraint, and respect for one another in the process of establishing and/or modifying their programs.
- Reciprocal access to one another's facilities should be the norm, and scientific exchange of data should be encouraged.
- Hardware cooperation should be based on mutual advantage and clear technical and management interfaces.
- Coordination among separate projects should be implemented whenever possible.
- Fair acknowledgment of the contribution of the other parties should be recognized and publicized.
- Upholding international agreements should have priority.

Some general rules

- The formulation of global road maps, involving the present set of space nations plus those which appear on the scene becomes everyday a pressing evidence
- However the technology transfer (ITAR) issue requires new approaches
- Some space agencies are not necessarily in full support of such an approach, which deprives them of their freedom and ability to keep a certain control
- However, can they maintain global control because of their own idiosyncrasies and national pride?
- A new system of governance must nevertheless be established particularly in Earth sciences



Future of International cooperation

- International cooperation in space science is a must
- It can only expand, including all space fairing nations and all talents from all nations
- Reciprocity between partners should be the rule
- Coordination of missions is easier to implement and should face no difficulties



Thanks

